



## Improved age control on early *Homo* fossils from the upper Burgi Member at Koobi Fora, Kenya



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### ABSTRACT

To address questions regarding the evolutionary origin, radiation and dispersal of the genus *Homo*, it is crucial to be able to place the occurrence of hominin fossils in a high-resolution chronological framework. The period around 2 Ma (millions of years ago) in eastern Africa is of particular interest as it is at this time that a more substantial fossil record of the genus *Homo* is first found. Here we combine magnetostratigraphy and strontium (Sr) isotope stratigraphy to improve age control on hominin-bearing upper Burgi (UBU) deposits in Areas 105 and 131 on the Karari Ridge in the eastern Turkana Basin (Kenya). We identify the base of the Olduvai subchron (bC2n) plus a short isolated interval of consistently normal polarity that we interpret to be the Pre-Olduvai event. Combined with precession-forced (~20 kyr [thousands of years]) wet–dry climate cycles resolved by Sr isotope ratios, the magnetostratigraphic data allow us to construct an age model for the UBU deposits. We provide detailed age constraints for 15 hominin fossils from Area 131, showing that key specimens such as cranium KNM-ER 1470, partial face KNM-ER 62000 and mandibles KNM-ER 1482, KNM-ER 1801, and KNM-ER 1802 can be constrained between  $1.945 \pm 0.004$  and  $2.058 \pm 0.034$  Ma, and thus older than previously estimated. The new ages are consistent with a temporal overlap of two species of early *Homo* that can be distinguished by their facial morphology. Further, our results show that in this time interval, hominins occurred throughout the wet–dry climate cycles, supporting the hypothesis that the lacustrine Turkana Basin was a refugium during regionally dry periods. By establishing the observed first appearance datum of a marine-derived stingray in UBU deposits at  $2.058 \pm 0.034$  Ma, we show that at this time the Turkana Basin was hydrographically connected to the Indian Ocean, facilitating dispersal of fauna between these areas. From a biogeographical perspective, we propose that the Indian Ocean coastal strip should be considered as a possible source area for one or more of the multiple *Homo* species in the Turkana Basin from over 2 Ma onwards.

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### Introduction

The time interval around 2 Ma (millions of years ago) is considered to be a key period in hominin evolutionary history in Africa. It marks the first appearance of several taxa in the genus *Homo* (Wood, 2009; Wood and Leakey, 2011; Leakey et al., 2012), an increase in hominin brain size (e.g., Elton et al., 2001; Potts, 2011),

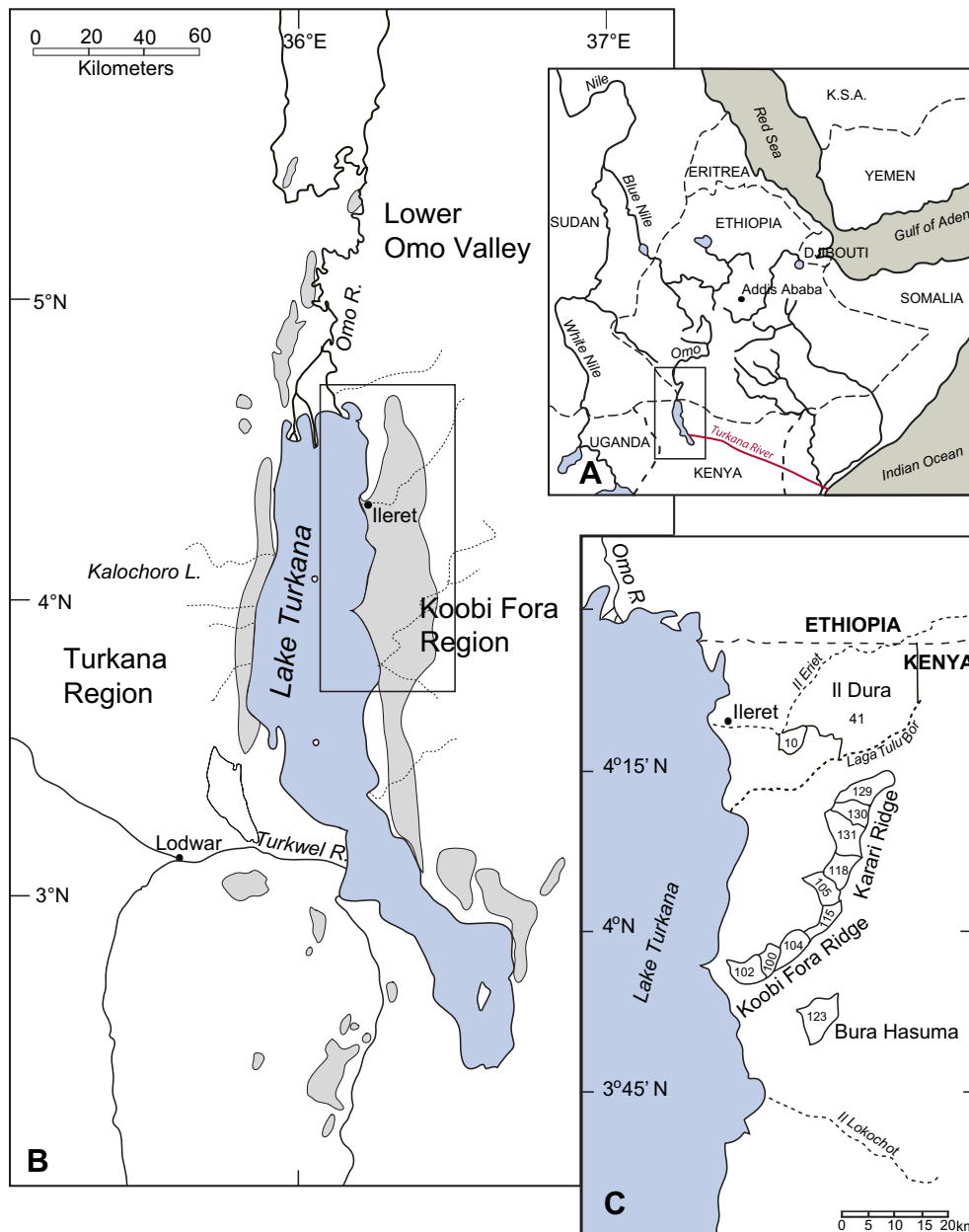
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and the earliest evidence for the use of stone tools to butcher both terrestrial and aquatic fauna (Braun et al., 2010). Moreover, at that time at least three and possibly more hominin species co-existed in the Turkana Basin (Kenya, Ethiopia; Fig. 1; e.g., Wood and Leakey, 2011; Leakey et al., 2012; Wood, 2012). Many issues pertinent to hominin origins, radiation and dispersal in this time interval remain unresolved, including the phylogenetic relationships between different broadly contemporaneous hominin species in Africa (Spoor et al., 2007; Berger et al., 2010; Pickering et al., 2011), the ecological niches occupied by coexisting hominin species (Ungar and Sponheimer, 2011), and the possible influence of climatic and environmental change on hominin dispersal and evolutionary history (Behrensmeyer, 2006; deMenocal, 2011). In order to address these issues, it is necessary to place hominin fossils

from this crucial time interval in a high-resolution chronological framework providing temporal linkage to local and global paleo-climate records as well as to well-dated archeological records.

Recently, Joordens et al. (2011) developed a novel climate proxy specifically for the Turkana Basin lacustrine deposits. The proxy method consists of combining strontium (Sr) isotope stratigraphy (measuring  $^{87}\text{Sr}/^{86}\text{Sr}$  of fish apatite sampled from lake sediments) with high-resolution magnetostratigraphy (providing the required age tie points) to resolve precession-forced wet–dry climate cycles. To obtain a high-resolution astronomically calibrated age and climate framework, we employ age models based on  $^{40}\text{Ar}/^{39}\text{Ar}$  ages of radiometrically-dated tuff layers (McDougall and Brown, 2006), as well as paleomagnetic reversal ages derived from astronomical calibration (Lourens et al., 2004; Gradstein et al., 2012). Hence, for



**Figure 1.** A. Location of the Turkana Basin in northeastern Africa. In red: postulated route of the 'Turkana River' connecting Turkana Basin and Indian Ocean around ~2 Ma (after Feibel, 1993). B. Present-day Lake Turkana with Pliocene and Pleistocene deposits indicated in grey. C. Location of a selection of the numbered paleontological collecting areas in the Koobi Fora region on the northeastern coast of Lake Turkana. Dotted lines indicate ephemeral rivers. (For interpretation of the references to colour in this figure legend, the reader is referred to the web version of this article.)

**Table 1**

Overview of ages of stratigraphic levels as used in this study, tied to the astronomical timescale (Lourens et al., 2004; Gradstein et al., 2012).

Name level	Radiometric age	Used in this study
KBS Tuff	1.869 ± 0.021 <sup>a</sup>	1.876 ± 0.021 <sup>b</sup>
bC2n		1.945 ± 0.004 <sup>c</sup>
Pre-Olduvai event		1.977 ± 0.002 <sup>d</sup>
Réunion-2 event	2.06–2.08 ± 0.01 <sup>e</sup>	2.082 ± 0.037 <sup>f</sup>

<sup>a</sup> Age provided by McDougall and Brown (2006) based on a Fish Canyon Tuff age of 28.11 Ma (Spell and McDougall, 2003).

<sup>b</sup> Age when normalized to the astronomically calibrated Fish Canyon Tuff age of 28.201 Ma (Kuiper et al., 2008).

<sup>c</sup> Horng et al. (2002).

<sup>d</sup> Channell and Guyodo (2004).

<sup>e</sup> Age of reversal obtained by stratigraphic scaling between radiometric ages based on a Fish Canyon Tuff age of 28.11 Ma (Kidane et al., 2007).

<sup>f</sup> Radiometric age of tuff with normal polarity from Gammari, Ethiopia (Kindley et al., 2012).

internal consistency in the present study, we adjust available <sup>40</sup>Ar/<sup>39</sup>Ar ages to an astronomical calibration (Kuiper et al., 2008). We apply this approach to date the fossil-rich hominin-bearing deposits of the ~2.00–1.87 Ma upper Burgi Member on the Karari Ridge in the eastern Turkana Basin (Fig. 1). Since the early 1970s, important fossils attributed to early *Homo* have been and continue to be found in these particular deposits (Feibel et al., 1989; Wood, 1991; Wood and Leakey, 2011; Leakey et al., 2012). Examples include the crania KNM-ER 1470 (Leakey, 1973a) and KNM-ER 62000 (Leakey et al., 2012), the mandible KNM-ER 1802 (Leakey, 1974) and the femur KNM-ER 1481 (Leakey, 1973a). This approach allows us to significantly improve age control on these and 11 other important hominin fossils from this region, and to place them in a high-resolution climatic and environmental context.

## Background

### Setting

The Turkana Basin is a hydrographic and sedimentary system encompassing about 131,000 km<sup>2</sup> of northern Kenya and southern Ethiopia (Fig. 1; Feibel, 2011). Since ~4.2 Ma, the paleohydrology of the Turkana Basin has been characterized by recurrent alternation of a major meandering river-floodplain system, and a lake system filling the basin. A persistent hydrological feature, the (paleo-)Omo River draining the vast Ethiopian Highlands has always supplied most (~90%) of the water in the basin (Feibel, 1999, 2011). Between ~2.0 and 1.5 Ma, the Turkana Basin held the long-lived, basin-wide Lake Lorenyang (Feibel, 2011), with a rich and diverse molluscan fauna demonstrating freshwater, non-alkaline conditions (Cerling, 1979; Vonhof et al., 2013). The presence of a hydrological connection between Lake Lorenyang and the Indian Ocean is documented by the first appearance of a marine-derived stingray in the basin at the onset of the Lorenyang lake phase. This fluvial system (named the ‘Turkana River’) exited the basin in the southeast and presumably followed the topography, controlled by the Central African Rift System, via the Anza Graben and Lamu Embayment to the Indian Ocean (Fig. 1; Feibel, 1993, 2011).

Our study focuses on the Karari Ridge in the eastern part of the basin (Fig. 1) where the early part of the Lorenyang lake phase is recorded in deposits of the upper Burgi Member (UBU; Koobi Fora Formation), consisting of pelagic (deep lake), fluctuating lake margin, and deltaic sediments (Brown and Feibel, 1986, 1991; Feibel et al., 1991). The base of the lacustrine-deltaic UBU consists of a diachronous erosional unconformity surface, which due to local tectonic processes is expected to be of increasing age on a south to north axis from the Koobi Fora Ridge to the Karari Ridge and the Ileret and Il Dura regions (see Fig. 1; Feibel, 2011).

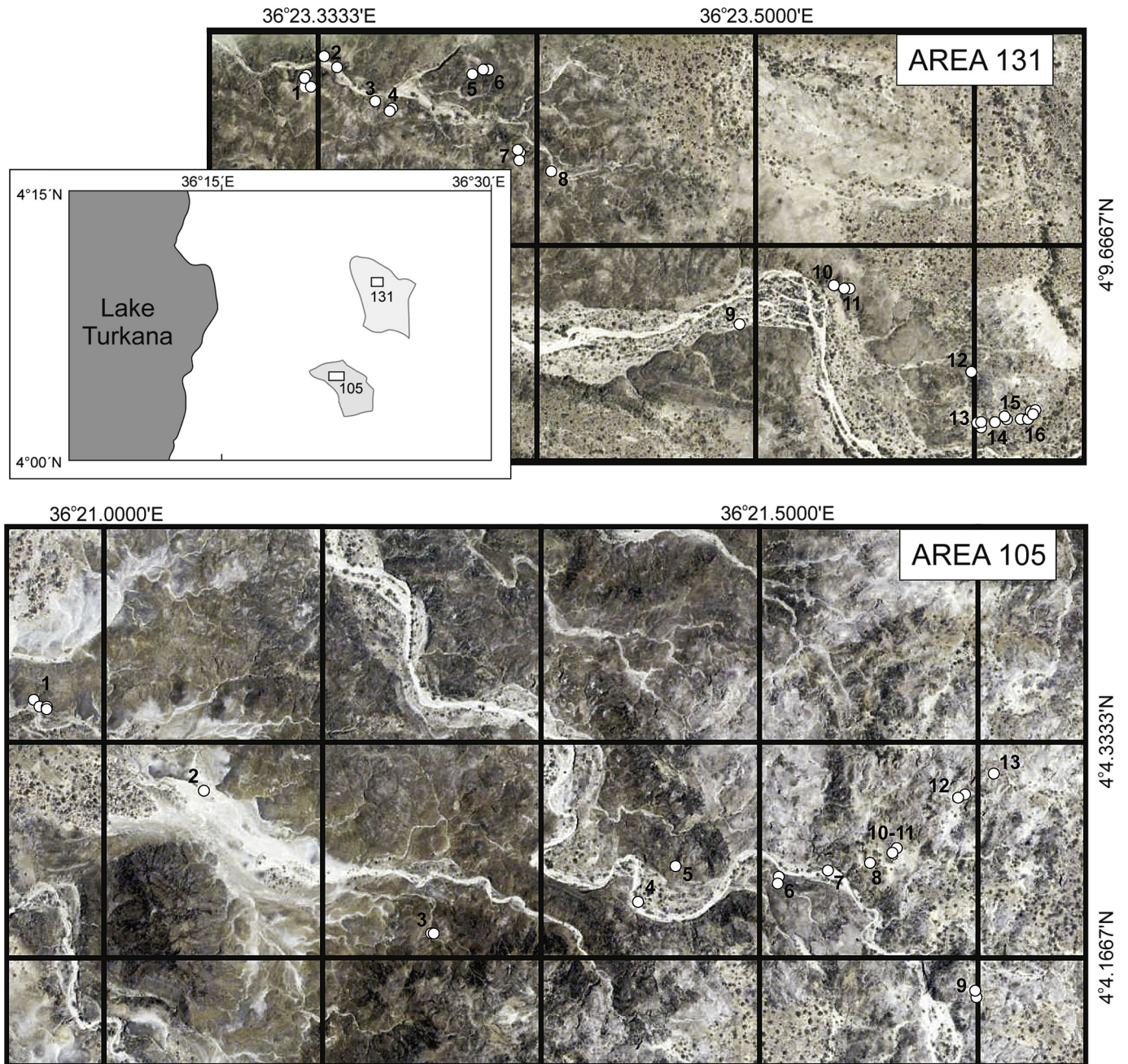
### Radiometric age control of early Lake Lorenyang deposits

Early Lake Lorenyang deposits equivalent to the UBU are upper Member G (~submember G-13 to Tuff H-2) of the Shungura Formation in the northern Turkana Basin, and the upper part of the Kalochoro Member (Nachukui Formation) in the western part of the Turkana Basin (Fig. 1; de Heinzelin, 1983; Lepre et al., 2011). Sedimentation in the northern and western Turkana Basin was continuous and hence in these regions the transition from fluvial-floodplain to lacustrine conditions (the onset of Lake Lorenyang) is documented in an uninterrupted sequence. In the eastern Turkana Basin, however, this was not the case, and an estimated ~400 kyr (thousands of years) time interval below the base of the lacustrine-deltaic UBU is missing (Brown and Feibel, 1986). The upper Burgi Member, Kalochoro Member, and Member G (plus submember H-1) are all capped by the KBS Tuff (see Table 1 for ages used in this study). Except for this well-established age of the top of the UBU, the age of UBU deposits on the Karari Ridge is at present poorly constrained. The deposits contain no additional tuff layers that could be radiometrically dated. Also, there are no regionally continuous marker beds that would allow correlation with the astronomically dated UBU deposits on the Koobi Fora Ridge (Joordens et al., 2011).

### Paleomagnetic age control of early Lake Lorenyang deposits

Previous paleomagnetic studies on early Lake Lorenyang sediments have been carried out in the northern Turkana Basin (Kidane et al., 2007), the western Turkana Basin (Lepre et al., 2011) and the eastern Turkana Basin (Brock and Isaac, 1974; Hillhouse et al., 1977; Braun et al., 2010; Joordens et al., 2011). In the northern Turkana Basin, Kidane et al. (2007) documented the presence of the base of the Olduvai chron (bC2n) dated at 1.945 ± 0.004 Ma (Horng et al., 2002; Lourens et al., 2004; Gradstein et al., 2012) and the presence of two normal polarity events below bC2n that were dated and interpreted by Kidane et al. (2007; Table 1) to belong to the Réunion subchron (Chamalaum and McDougall, 1966). The younger event was at about the same level as the onset of the Lake Lorenyang deposits (de Heinzelin, 1983; Kidane et al., 2007). In the western Turkana Basin, Lepre et al. (2011) documented the base of the Olduvai chron, plus a short normal polarity event ~30 m below bC2n at ~2 m above the onset of Lake Lorenyang deposits. These paleomagnetic data from the northern and western Turkana Basin show that Lake Lorenyang started at or just before the occurrence of a short paleomagnetic normal event, likely – in view of its short duration – the younger ‘second Réunion’ event, also named RU-2 excursion (e.g., Kidane et al., 1999; Quidelleur et al., 2010).

The age and especially the ubiquity of this RU-2 excursion are still debated (Carlut et al., 1999; Kidane et al., 1999, 2007; Lanphere et al., 2002; Quidelleur et al., 2010; Ellis et al., 2012; Kindley et al., 2012). Globally distributed volcanic and sedimentary records show that the older Réunion event (RU-1) is associated with a large dipole intensity decrease and is recorded in many sequences (Quidelleur et al., 2010). The dipole intensity decrease was not as pronounced during RU-2 and consequently, RU-2 is not present as a full directional change in many sections worldwide, but rather appears in certain places as a geomagnetic excursion during an episode of increased secular variation (Quidelleur et al., 2010). Lanphere et al. (2002) proposed, based on correlation of RU-2 with the Huckleberry Ridge Tuff from Yellowstone (USA), that RU-2 be named the Huckleberry Ridge event. However, recent dating of the Huckleberry Ridge tuff complex by Ellis et al. (2012) has cast doubt on this association so we prefer to indicate this excursion by the name ‘second Réunion’ or RU-2. A normal polarity excursion of similar age as RU-2 (in some cases accompanied by another one

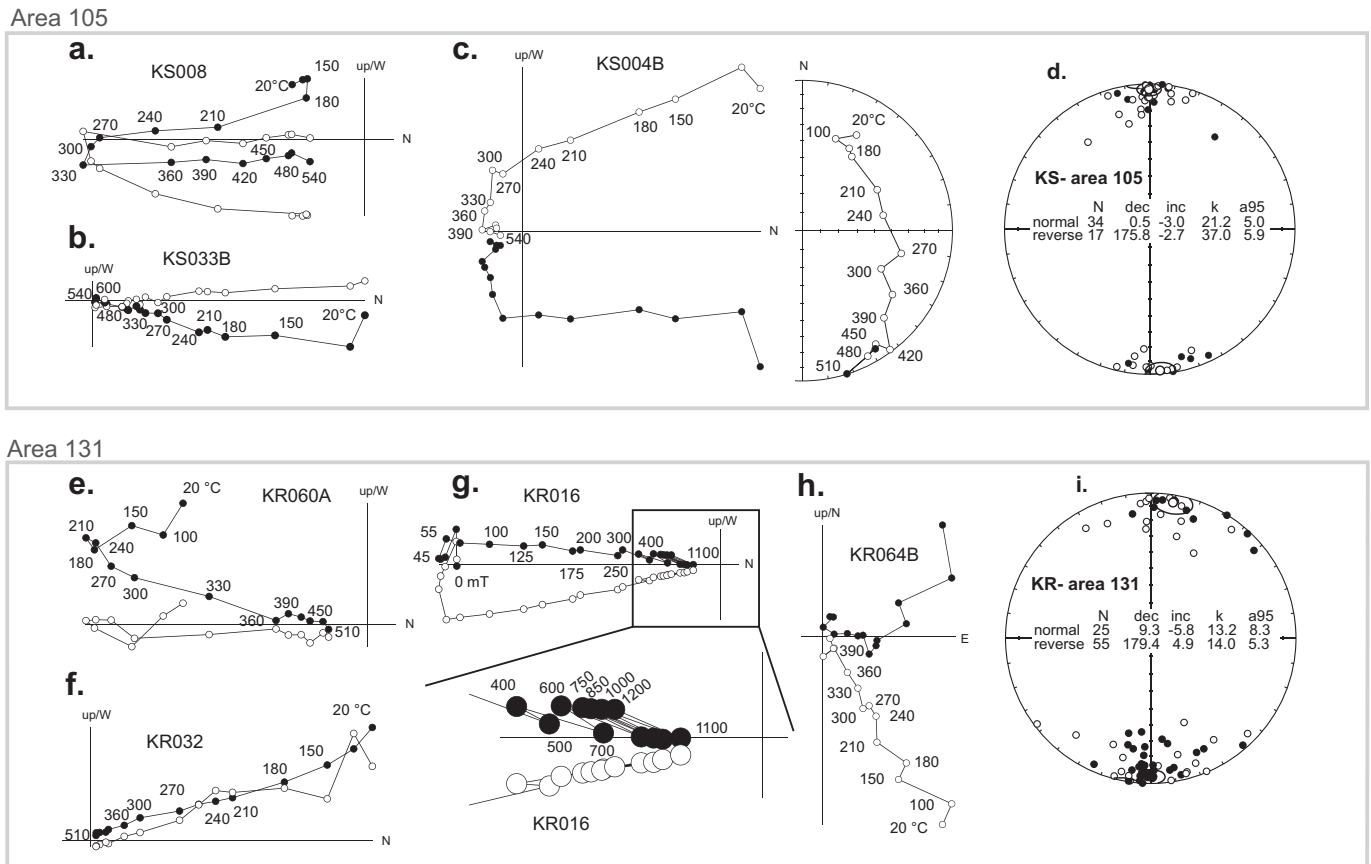


**Figure 2.** Map of sample locations within Koobi Fora Areas 131 and 105, and Google Earth images with detailed locations of sampled subsections (white dots) in Areas 105 and 131. Detailed information on the subsections is provided in [Tables A.1 and A.2 in the Appendix](#).

with a similar age as RU-1) has been detected in several regions in the eastern part of Africa: the Afar depression in Ethiopia (Kidane et al., 1999), the northern Turkana Basin in Ethiopia (Kidane et al., 2007) and western Turkana Basin (Lepre et al., 2011) in Kenya, and Sterkfontein in South Africa (Dirks et al., 2010; Herries and Shaw, 2011). Kindley et al. (2012) confirmed previous work by Kidane et al. (1999) that RU-2 was recorded in Gammari in Ethiopia (Table 1). This indicates that in eastern Africa, there is evidence for the occurrence of a double Réunion subchron. We therefore conclude that the RU-2 can be used as a local time anchor marking the onset time of Lake Lorenyang in the Turkana Basin. We assume that at present  $2.08 \pm 0.04$  Ma (Kindley et al., 2012) is the best approximation for the age of RU-2 (representing the minimum age

of the onset of Lake Lorenyang) and hence for the maximum age of the UBU sections in Areas 131 and 105.

In the eastern Turkana Basin (Area 41; Fig. 1), Braun et al. (2010) documented the base of the Olduvai chron at  $\sim 14$  m below the KBS Tuff, plus a short normal polarity excursion at  $\sim 5$  m below the base of the Olduvai chron. Joordens et al. (2011) found in Area 102 (Fig. 1) the base of the Olduvai chron at  $\sim 80$  m below the KBS Tuff. At the Karari Ridge, our study area, the available paleomagnetic data are equivocal. The results of Brock and Isaac (1974) in Area 105 on the Karari Ridge (Fig. 1) suggest the presence of the base of the Olduvai chron  $\sim 20$  m below the KBS Tuff, based on exclusively normal directions above that level and a mix of reversed and normal polarity directions below. In Area 131 on the Karari Ridge (Fig. 1), early



**Figure 3.** Paleomagnetic results. a–c and e–h Representative demagnetization behaviors observed in the section of Areas 105 (KS section) and 131 (KR section) respectively. Solid (open) points are projections on the horizontal (vertical) plane. d and i. Stereographic projections of ChRM directions. Solid (open) points are projections on the lower (upper) hemisphere. Mean of normal and reversed directions are indicated by larger points and given in central table (dec, declination; inc, inclination; k, precision parameter;  $\alpha_{95}$ , 95% confidence cone). c. Typical thermal demagnetization path (steps given in °C) going to reversed direction treated by great circle analysis (McFadden and McElhinny, 1988). g. Typical AF demagnetization with biasing above 200 Oersted (see enlarged inset) interpreted to result from gyroremanence acquisition during AF treatment. h. Widely outlying direction interpreted as transitional.

work of Hillhouse et al. (1977) suggests that normal paleomagnetic polarity extends to ~14–15 m below the KBS Tuff, followed by one sample with reversed polarity, then normal polarity again between 17 and 20 m below the KBS Tuff, followed by reversed polarity to ~35 m below KBS. In the lowermost 5 m of the UBU section they find normal polarity again (Hillhouse et al., 1977). McDougall et al. (2012) mention that the polarity transition at ~15 m below the KBS Tuff in Area 131 almost certainly represents the bottom of the Olduvai chron, whereas the piece of section with normal polarity lower in the sequence may either be spurious (normal polarity overprint) or represents a geomagnetic excursion. The relatively low sample resolution in these studies from the 1970s, and the possible occurrence of normal overprint, precludes the precise definition of polarity reversal positions necessary for reliable magnetostratigraphic age constraints. This lack of reliable paleomagnetic results in the lower part of the UBU sections on the Karari Ridge has motivated the high-resolution magnetostratigraphic analysis presented here.

## Magnetostratigraphy

### Methods

Paleomagnetic sampling was performed in 2010 in sediments of the upper Burgi Member in Areas 105 and 131 on the Karari Ridge in the eastern Turkana Basin (Fig. 1). We obtained a complete

stratigraphic succession undisturbed by faults, by stacking several carefully correlated nearby subsections (Fig. 2, Tables A.1 and A.2 in the Appendix). This was done by physically tracing distinctive correlative beds from one subsection to the next. Subsections were measured and trenched, and the lithology was described, after which paleomagnetic samples were taken at ~50 cm intervals. Paleomagnetic samples consisted of standard 2.5 cm-diameter cylindrical rock cores drilled with an electric drill mounted with a diamond-coated bit and cooled with an electric air compressor, both devices powered by a portable generator. Loose sandy intervals were sampled by pressing 2.5 cm-diameter quartz glass cylinders, filled with a sodium silicate solution for in situ hardening, into the sediment. Samples (cores and cups) were oriented with a compass and clinometer and no dip corrections were necessary for these flat lying sediments. Orientations were corrected for the local 1.5° declination. Paleomagnetic processing was performed using 2G DC squid magnetometers at the paleomagnetic laboratories of the Rennes 1 University (France) and the CENIEH in Burgos (Spain). Demagnetization procedures followed methods described in Dupont-Nivet et al. (2008) and Joordens et al. (2011). Detailed analysis of pilot samples for both studied areas revealed the same type of demagnetization behaviors, very similar to those obtained from Area 102 presented in Joordens et al. (2011). Alternating Field (af) demagnetization of pilot samples was often biased by gyroremanence at higher applied fields while thermal demagnetization gave superior Characteristic Remanence Magnetization

**Table 2**  
Characteristic Remanent Magnetization (ChRM) directions from Area 105 (section KS).

ID#	Level (m)	Dec. (°)	Inc. (°)	MAD (°)	VGP (°)	Q	Tinf (°C)	Tsup (°C)	Dir.
KS063	31.8	355.9	-22.1	10.5	74.0	1	150	570	Free
KS062	31.5	1.9	-8.6	14.9	81.5	1	240	420	Free
KS061	30.9	66.7	-3.5	8.2	23.1	3	180	600	Free
KS060B	30.3	356.8	-9.6	13.5	50.3	1	210	420	Free
KS060A	30.3	324.6	-27.4	4.2	80.6	1	210	450	Free
KS059A	29.8	358.5	-1.2	19.3	85.2	1	240	420	Free
KS057	28.0	358.3	-7.3	16.0	82.1	1	180	400	Free
KS056	27.9	9.5	-12.0	17.4	76.2	1	150	510	Free
KS055	27.5	348.9	-27.0	9.8	68.7	1	180	570	Free
KS054	26.7	359.1	-10.2	9.6	80.8	1	180	570	Free
KS052 <sup>af</sup>	23.5	357.6	5.6	11.4	87.3	1	100	500	Free
KS051A	23.2	0.9	-6.0	25.5	82.9	1	180	480	Free
KS049	22.7	359.2	18.9	20.3	84.2	1	210	510	Free
KS048	22.5	352.5	-11.1	6.8	77.8	1	180	450	Free
KS042	21.5	357.9	-12.4	18.2	76.5	1	240	400	Free
KS044 <sup>af</sup>	21.5	346.6	4.4	1.4	79.5	1	45	600	Free
KS040	20.5	342.6	-4.4	5.8	71.5	1	210	510	Free
KS039	20.4	359.2	9.1	10.6	89.0	1	180	420	Free
KS038A	19.4	355.8	-18.4	10.0	75.9	1	210	360	Free
KS037A	18.5	359.8	-10.1	13.4	80.9	1	180	540	Free
KS036A <sup>af</sup>	18.0	0.1	0.8	2.1	86.4	1	75	500	Free
KS034A	16.5	17.6	-2.1	6.7	71.7	1	210	450	Free
KS033B	16.0	15.6	-8.6	6.6	72.3	1	240	400	Free
KS031A	15.4	7.4	2.6	11.3	82.1	1	240	450	Free
KS032	15.1	345.0	-19.3	4.8	69.6	1	180	450	Free
KS030A	14.6	358.2	0.7	11.7	85.9	1	240	450	Free
KS028	13.3	2.0	13.6	4.1	86.5	1	75	600	Free
KS027	12.7	3.4	-1.4	6.2	84.2	1	180	510	Free
KS026	12.0	7.7	-1.2	13.1	81.0	1	270	540	Free
KS025	11.7	2.0	4.5	12.0	87.3	1	240	540	Free
KS024	11.4	35.2	23.3	11.3	54.2	1	20	450	Free
KS023	10.9	9.0	-6.3	5.7	78.5	1	210	480	Free
KS022	10.1	5.0	0.0	12.4	83.6	1	240	400	Free
KS014A	10.0	350.4	10.3	12.9	80.4	1	240	420	Free
KS015B	9.7	2.9	-2.2	5.0	84.1	2	150	420	Forced
KS016A	9.1	356.4	21.6	11.3	-83.8	3	240	380	Free
KS016B	9.1	182.8	-18.5	9.8	82.0	2	210	450	gc
KS017B	8.8	5.7	8.6	5.4	84.3	3	0	340	Free
KS018B	8.5	180.5	-5.5	25.5	-88.7	2	100	480	gc
KS019B	7.9	186.8	-12.9	23.3	-82.8	2	180	450	gc
KS020A	7.5	181.4	-5.5	18.8	-88.1	2	0	340	gc
KS020B	7.5	35.5	10.3	7.2	54.6	3	210	360	Free
KS021	6.9	187.0	-6.2	19.3	-83.0	2	100	480	gc
KS013	6.4	170.7	0.6	9.1	-79.8	1	240	570	Free
KS012	6.1	189.0	7.2	7.1	-78.2	1	330	510	Free
KS011	5.5	162.4	7.1	11.4	-70.9	1	210	450	Free
KS010	5.1	173.0	-2.3	14.7	-82.4	1	240	570	Free
KS009A	4.4	161.0	-12.7	11.5	-70.9	1	210	540	Free
KS008	4.3	180.7	3.7	5.7	-84.1	1	330	510	Free
KS006B	3.2	155.0	4.7	23.2	-64.2	2	270	38	Forced
KS005A	2.7	168.8	-3.5	22.4	-78.6	2	350	450	Forced
KS004B	1.9	166.0	10.4	12.1	-73.2	2	100	450	gc
KS003A	1.4	170.8	-2.9	18.5	-80.5	2	280	510	Forced
KS002	0.8	179.5	-6.5	16.9	-89.1	2	270	310	Forced
KS001	0.1	193.4	-2.9	15.8	-76.4	1	300	510	Free
Averages	N	Dec. (°)	Inc. (°)	$\alpha_{95}$ (°)	Tinf (°C)	Tsup (°C)			
Mean N	34	0.5	3.0	5.0	193	463			
Mean R	17	175.8	-2.7	5.9					

ID #, Sample identification; af, Alternating field demagnetization; Level, Stratigraphic position in meters (m) with respect to base of section; Dec., Declination; Inc., Inclination; MAD, Maximum Angular Deviation; VGP, Virtual Geomagnetic Pole latitude; Q, Quality factor (1, reliable direction; 2, reliable polarity; 3, unreliable polarity); Tinf, Lower demagnetization temperature of ChRM; Tsup, Higher demagnetization temperature of ChRM; Dir., Direction fitting analysis (free, free line fit; forced, forced line fit; gc, great circle). Averages, Mean of normal polarity (Mean N) and reversed polarity (Mean R) ChRM directions;  $\alpha_{95}$ , Angular radius of 95% confidence interval; Tinf, Mean lower demagnetization temperature; Tsup, Mean higher demagnetization temperature.

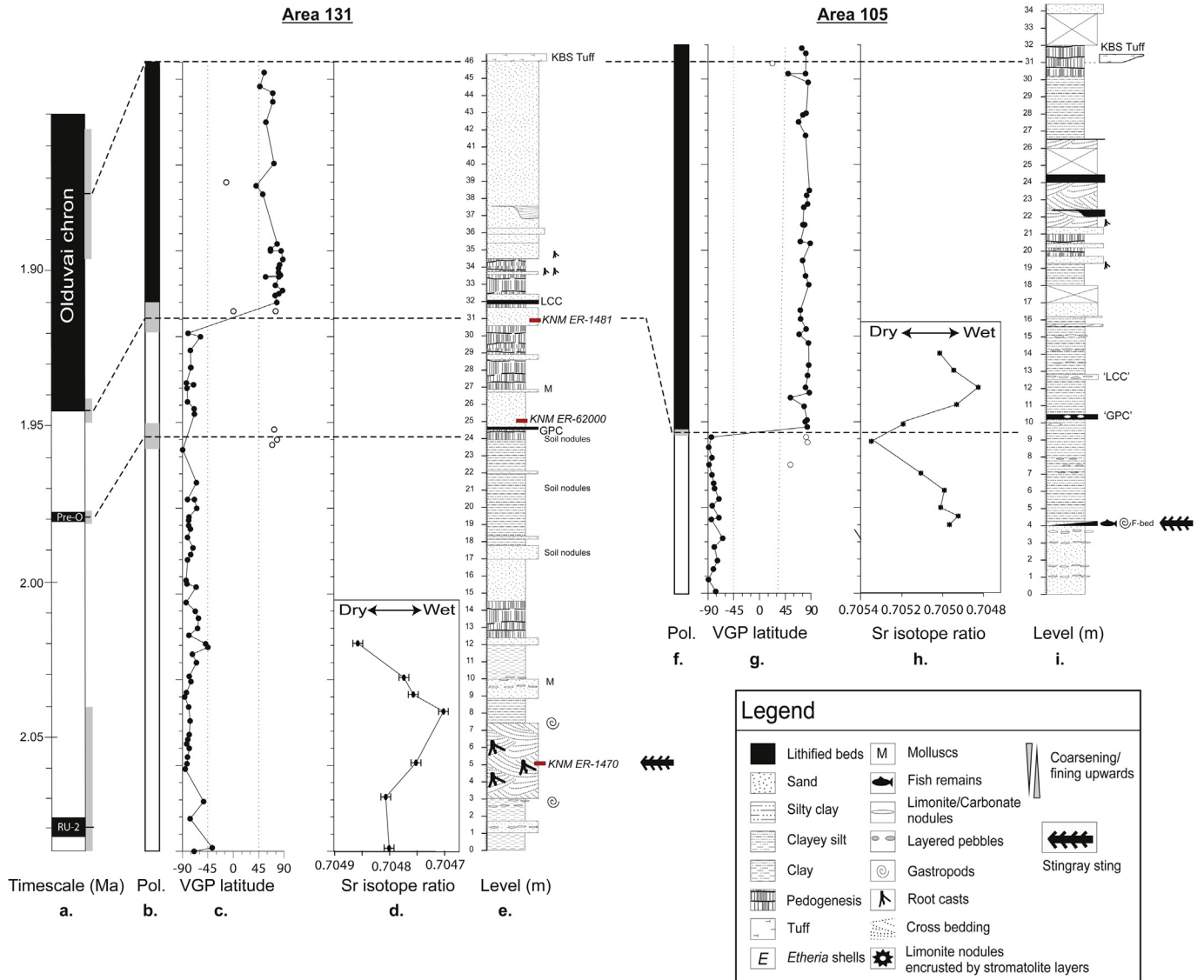
(ChRM) component separation and was used in all following processing (Fig. 3). Of all 153 demagnetized samples, only four with erratic paths and weak magnetization yielded no interpretable paleomagnetic directions. For most other demagnetized samples (91 out of 149; Tables 2 and 3), after removal of a common normal overprint up to 200–300 °C, the ChRM was determined by a steady decay toward the origin up to 400–550 °C (Fig. 3), typical of Ti-rich titanomagnetites carrying the remanence similarly to results from Area 102 (Joordens et al., 2011). Characteristic Remanence Magnetization directions were then calculated using least square analysis (Kirschvink, 1980) on a minimum of four consecutive steps of the ChRM. These line fits provide reliable paleomagnetic direction estimates (labeled 'quality 1') and were obtained without forcing to the origin (Tables 2 and 3). For most other samples (38), obtained directions provide reliable polarity estimates (labeled 'quality 2') but directions were obtained using other methods for the following reasons: (1) noisy demagnetization paths required forcing line fits through the origin, or (2) a normal overprint partially overlapping with the reversed ChRM was clearly indicated by demagnetization paths along a great circle on stereographic projections (Fig. 3c). In the latter cases, best fit great circle analyses were performed using the mean of previously obtained 'quality 1' reversed directions as set points (McFadden and McElhinny, 1988). Finally, 10 samples were qualified unreliable in direction and polarity ('quality 3') because they yielded widely aberrant directions or isolated levels of opposite polarity. These directions are interpreted either as transitional during reversals and excursions, or as delayed acquisition in sediments below a reversal. The obtained ChRM directions have maximum angular deviation (MAD) below 30° (average is 12°). Excluding quality 3 directions, they cluster around antipodal normal- and reverse-polarity mean directions (Fig. 3d and i) and pass the reversals test (McFadden and McElhinny, 1990) with class C. This suggests stability of the ChRMs, successful exclusion of normal overprint bias and a reliable magnetic polarity zonation for the sections in both Areas 105 (KS sections) and 131 (KR sections).

## Results

**Base of the Olduvai chron in Areas 105 and 131** The transition of reversed polarity directions in the lower part of the UBU, to normal polarity in the upper part of the UBU, unequivocally and precisely locates the base of the Olduvai chron in the sampled sections (Fig. 4, Tables 2 and 3). In the UBU section of Area 105, the reversal to bC2n is precisely indicated between 21.3 and 21.9 m below the KBS Tuff (Fig. 4a, f and g). Three isolated directions of seemingly normal polarity, positioned just below and up to 2.0 m below the reversal, may be interpreted as resulting from delayed acquisition of a normal overprint (Van Hoof and Langereis, 1991). In the UBU section of Area 131, the base of the Olduvai chron is situated between the 30.2 and 32.0 m levels (14.0–15.8 m below the KBS Tuff; Fig. 4a–c). In between, two samples taken in sands at the 31.5 m level yield mixed results of transitional and normal directions, respectively, suggesting recording of the polarity reversal (Table 2). These results confirm the results of earlier studies that located the bottom of the Olduvai chron in Area 105 at ~20 m below KBS Tuff (Brock and Isaac, 1974) and in Area 131 at ~15 m below KBS Tuff (Hillhouse et al., 1977).

**Pre-Olduvai event in Area 131** In Area 131, a short isolated interval of consistently normal polarity directions is situated within the 23.7–24.6 m level, at 21.4–22.3 m below KBS Tuff (Fig. 4a–c, Table 3). This may result from normal overprinting coinciding with soil formation, however, such overprinting has not been found in any of the other soil layers in these sequences. Delayed acquisition is also unlikely given the large stratigraphic interval (~7 m) below the normal





**Figure 4.** Overview of the paleomagnetic and Sr isotope results from Areas 131 and 105. a. Timescale with periods of normal polarity in black and reversed polarity in white. Pre-O: Pre-Olduvai event. RU-2: second Réunion event. In grey: error bars on the ages of KBS Tuff and paleomagnetic reversals (Table 1). b. Magnetic polarity in Area 131. c. Virtual Geomagnetic Pole (VGP) latitude (in degrees) in Area 131. d. Sr isotope ratios in Area 131. e. Lithostratigraphy of the composite section in Area 131, with stratigraphic positions of hominin fossils KNM-ER 1481, KNM-ER 62000 and KNM-ER 1470 (see also Figure 5 and Table 7). f. Magnetic polarity in Area 105. g. VGP latitude (in degrees) in Area 105. h. Sr isotope ratios in Area 105. i. Lithostratigraphy of the composite section in Area 105. 'GPC' and 'LCC' indicates that these marker beds are not time equivalent with the GPC and LCC beds in Area 131. Note that the F-bed in Area 105 is not the same as marker bed F in Area 131 (Table 7).

catchment geology of the Omo and southwest drainage areas differs, the rivers have different chemical compositions and variation of relative river contribution to the lake causes changes in lake water chemistry (Yuretech, 1979). The (paleo-)Omo River catchment drains mainly Tertiary flood basalts with relatively low  $^{87}\text{Sr}/^{86}\text{Sr}$  values ( $\sim 0.704$ ), whereas river catchments in the southwest drain metamorphic Precambrian basement terrains with relatively high  $^{87}\text{Sr}/^{86}\text{Sr}$  values ( $\sim 0.720$ ; Talbot et al., 2000). Thus, during wet periods of strong monsoonal rainfall in the Ethiopian Highlands, the proportion of Omo water in the lake is expected to be relatively high, and  $^{87}\text{Sr}/^{86}\text{Sr}$  values of lake water would tend toward that of Ethiopian flood basalts ( $\sim 0.704$ ). During dry periods in Ethiopia, the proportion of Omo water in the lake would be somewhat lower and  $^{87}\text{Sr}/^{86}\text{Sr}$  values of lake water would increase.

Lacustrine Sr isotope ( $^{87}\text{Sr}/^{86}\text{Sr}$ ) variation is solely driven by water provenance, so variation in lake water  $^{87}\text{Sr}/^{86}\text{Sr}$  over time can be used as a proxy for variation in rainfall intensity over the Ethiopian Highlands. This variation is faithfully recorded in organisms

living in the water:  $^{87}\text{Sr}/^{86}\text{Sr}$  values of well-preserved aquatic fossils (e.g., fish bones and teeth) from lake sediments directly reflect  $^{87}\text{Sr}/^{86}\text{Sr}$  values of the host water, and are not affected by temperature variation or ecological parameters (Talbot et al., 2000; Vonhof et al., 2003). Due to the hydrologically conservative behavior of Sr in a lake system, Sr is sensitive to changes on orbital timescales but insensitive to changes on submillennial timescales: spectral analysis has shown that Sr isotope ratios reflect changes in monsoonal rainfall on precession ( $\sim 20$  kyr) timescales (Joordens et al., 2011).

#### Methods

Samples of unweathered sediment were taken at 1 m intervals in the same UBU subsections in Areas 105 and 131 where paleomagnetic samples were taken. Sediment samples (ca. 5 cm<sup>3</sup> each) were wet-sieved over a 600- $\mu\text{m}$  sieve and, when present, optically well-preserved fish fossils (bones, teeth) were selected using a binocular microscope. Excellent preservation of fish fossils was confirmed by



**Table 4**  
Sr isotope data of Area 105.

Sample code	Sample type	Strat. Level (m)	$^{87}\text{Sr}/^{86}\text{Sr}$	Standard error ( $2\sigma$ )
10KS-14	Fish bone	14.0	0.705015	0.000009
<i>K86-2380</i>	<i>Stingray sting</i>	<i>13.0</i>	<i>0.704946</i>	<i>0.000008</i>
<i>K86-2381</i>	<i>Stingray sting</i>	<i>12.0</i>	<i>0.704826</i>	<i>0.000008</i>
<i>K86-2397</i>	<i>Fish tooth</i>	<i>11.0</i>	<i>0.704932</i>	<i>0.000008</i>
10KS-10	Fish bone	10.0	0.705194	0.000011
10KS-9	Fish bone	9.0	0.705350	0.000015
10KS-7	Fish tooth	7.0	0.705107	0.000010
10KS-6	Fish bone	6.0	0.704992	0.000010
10KS-5	Fish bone	5.0	0.705010	0.000009
<i>K86-2382+</i>	<i>Stingray sting</i>	<i>4.5</i>	<i>0.704924</i>	<i>0.000008</i>
<i>K86-2382</i>	<i>Stingray sting</i>	<i>4.0</i>	<i>0.704967</i>	<i>0.000009</i>

Sample code in italics: samples that were collected by CSF in 1986, values previously published in Table S5 of Joordens et al. (2011).

SEM inspection. Only minor diagenetic  $\text{CaCO}_3$  overgrowths were encountered in cracks or voids in some of the fish fossils. To effectively remove such possibly contaminant  $\text{CaCO}_3$ , fish fossil fragments were leached for ~1 h in 5 N acetic acid prior to Sr separation. The remaining fish apatite was dissolved in 3 N  $\text{HNO}_3$ , after which Sr was separated using Elchrom Sr spec ion exchange resin.  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios were analyzed on a Finnigan MAT 262 mass spectrometer at the department of Earth Sciences of the VU University Amsterdam, running a triple jump routine and applying exponential fractionation correction. All  $^{87}\text{Sr}/^{86}\text{Sr}$  data are reported relative to a value of 0.710243 for the routinely analyzed NBS 987 standard. Strontium in blanks was <0.1% of the total Sr in the samples.

## Results

From the 31 sediment samples taken in Area 105, only six yielded fish apatite suitable for Sr isotope analysis. In Area 131, 46 sediment samples were taken and 11 yielded sufficient and suitable apatite. This can be explained by the fact that in the largely deltaic UBU deposits on the Karari Ridge, with large variations in depositional environments, fish apatite is not as consistently abundant and well-preserved throughout the successions as in more pelagic sediments such as the UBU in Area 102 (Joordens et al., 2011). To complement the datasets, we included six Sr isotope values measured on fish apatite previously sampled by one of us (CSF) from well-constrained stratigraphic levels in the same sections in Areas 105 and 131 (Tables 4 and 5). The Sr isotope records from these Karari sequences are valuable because they contain climate information and give a measure of time: as shown by Joordens et al. (2011), a strontium isotope cycle recorded in lacustrine Turkana Basin deposits reflects a precession cycle of ~17–23 kyr duration. Moreover, the Sr isotope records can be directly correlated to both paleomagnetic reversals and hominin fossil occurrences.

**Sr isotope cycle in Area 105** Despite the difficulties in obtaining continuous suitable fish apatite records, we were able to capture a single well-resolved Sr isotope precessional cycle in ca. 6–7 m of section in the lower part of the UBU sequence in Area 105 (Fig. 4h, Table 4). The  $^{87}\text{Sr}/^{86}\text{Sr}$  values of the cycle in Area 105 range between 0.7054 and 0.7048; on average 0.7051, with amplitude ca. 0.0006. The Sr isotope cycle spans the base of the Olduvai chron (Fig. 4g and h), and thus can be compared with the Sr isotope cycle of the same age recorded in the UBU of Area 102 (Joordens et al., 2011). In both Areas 105 and 102, it was found that bc2n is situated close to a peak dry period in the Sr isotope cycle (in the phase going from dry to wet conditions), confirming the hypothesis that the observed cycles in these records are contemporaneous. The climate phase relationship appears to contradict that of the astronomically calibrated age of bc2n ( $1.945 \pm 0.004$  Ma; Horng

**Table 5**  
Sr isotope data of Area 131.

Sample code	Sample type	Level (m)	$^{87}\text{Sr}/^{86}\text{Sr}$	Standard error ( $2\sigma$ )
KJ10-10	<i>Sindacharax</i> tooth	33.5	0.705022	0.000011
10KR-29	Fish bones	29.0	0.704979	0.000008
10KR-22	Fish tooth	22.0	0.704966	0.000010
10KR-20	Fish tooth	20.0	0.704942	0.000008
10KR-12	<i>Distichodus</i> tooth	12.0	0.704872	0.000012
10KR-10	Fish bones	10.0	0.704788	0.000008
10KR-9	<i>Distichodus</i> tooth	9.0	0.704771	0.000009
10KR-8	Fish bones	8.0	0.704716	0.000008
<i>K86-2278</i>	<i>Stingray spine</i>	<i>5.0</i>	<i>0.704759</i>	<i>0.000009</i>
10KR-5	Fish bones	5.0	0.704766	0.000009
10KR-3	Fish bones	3.0	0.704821	0.000009
10KR-0	Fish bones	0.0	0.704815	0.000009

Sample code in italics: sample that was collected by CSF in 1986, value previously published in Table S5 of Joordens et al. (2011).

et al., 2002), which corresponds to a precession minimum and insolation maximum (Laskar et al., 2004) and hence to peak wet conditions (Tuenter et al., 2003). However, we point out that the error in this age of bc2n does not include the additional uncertainties (likely less than a precession cycle) that arise from the astronomical tuning process (Horng et al., 2002), and thus at present this age of bc2n cannot be used to resolve a climate phase relationship.

**Sr isotope cycle in Area 131** In Area 131, there appears to be a Sr isotope cycle in the lower part of the section, although this cycle is not as well-resolved as the cycle in Area 105 and may record (part of) a second cycle as well (Fig. 4d, Table 5). The  $^{87}\text{Sr}/^{86}\text{Sr}$  values of the cycle range between 0.7049 and 0.7047; on average 0.7048 with amplitude ca. 0.0002. Four additional Sr isotope values were obtained from fish apatite found at levels higher in the sequence of Area 131 (from above the 20 m level), but the number and resolution of these data points were too low to resolve more cycles. The Sr isotope values of these four samples are higher (0.7050) than those low in the section (Table 5). This suggests that the Sr isotope values in Area 131 increase over time, similar to the trend observed in the Sr isotope record of the UBU Area 102 where the values increase from ca. 0.7049 low in the section (below bc2n) to ca. 0.7055 higher in the section just below the KBS Tuff. As proposed for Area 102 (Joordens et al., 2011), the trend could be explained either by slowly changing water chemistry (e.g., due to a gradual change in river course, delta progradation) in the catchments feeding the Turkana Basin, or by a gradual relative decrease in monsoon rainfall in the Omo River catchment.

**Comparison of the Sr isotope values from Areas 105 and 131** The average value of the Sr isotope cycle in Area 131 (0.7048) is low compared with that of the cycle in Area 105 (0.7051). This is, in view of the apparent general trend in the Sr isotope record of the UBU, in agreement with the magnetostratigraphic results indicating that the cycle in Area 131 is significantly older than that of Area 105 (Fig. 4). In addition, the amplitude of the cycle in Area 131 (0.0002) appears to be considerably smaller than the amplitude of the cycle in Area 105 (0.0006), which suggests that the difference in monsoonal intensity between wet and dry phases of the precession cycle was smaller in this time period than around the base of the Olduvai chron.

## Age model calculations

### Methods

Our magnetostratigraphic results provide a well-dated temporal anchor to be employed in the calculation of age models for the UBU deposits on the Karari Ridge: the base of the Olduvai chron (bc2n,  $1.945 \pm 0.004$  Ma). The age of the younger Réunion event (RU-2;

2.08 ± 0.04 Ma) represents the minimum age of the onset of Lake Lorenyang, and hence the maximum age of the UBU sections in Areas 131 and 105. These two results are then used as a complement to the radiometrically-dated KBS Tuff (1.876 ± 0.021; Table 1) at the top of the sequence.

## Results

**Age model Area 105** Applying the available astronomically calibrated age tie points implies that in Area 105 the mean sediment accumulation rate between bC2n and the KBS Tuff is 31.3 ± 9 cm/kyr (Table 6). This value agrees well with the sediment accumulation rate of 31 cm/kyr that can be derived from the Sr isotope cycle in Area 105 (Fig. 4h) with ~6.5 m of sediments representing one precession cycle of 21 kyr. Linear extrapolation of the rate 31.3 ± 9 cm/kyr downwards from the KBS Tuff places the base of the section in Area 105 at 1.986 ± 0.052 Ma, taking into account the error of the age of the KBS Tuff and that of the sediment accumulation rate. The same extrapolation applied downwards from bC2n yields an age of 1.979 ± 0.014 Ma for the base of the section. This is in line with the documented absence of normal polarities below bC2n (Fig. 4f and g), indicating that the base of the UBU section in this area is slightly younger than the Pre-Olduvai event at 1.977 ± 0.002 Ma. It also suggests that the Pre-Olduvai event may be located just below the measured section in Area 105, and that future magnetostratigraphic analysis may find it there by extending the sampled section downward a few meters (Table 6).

**Age model Area 131** In Area 131, the mean sediment accumulation rate between bC2n and the KBS Tuff is 21.6 ± 7 cm/kyr (Table 6). Linear extrapolation of this rate downwards from bC2n to the stratigraphic level of the short normal polarity excursion below bC2n (m level 24; Fig. 4) yields an age of 1.981 ± 0.016 Ma for this level, which supports our interpretation that this polarity excursion is the Pre-Olduvai event dated to 1.977 ± 0.002 Ma. Linear extrapolation of the rate of 21.6 ± 7 cm/kyr downwards from the KBS Tuff to the base of the section yields an age of 2.114 ± 0.098 Ma for the base of the section, taking into account the error of the age of the KBS Tuff and the error of the sediment accumulation rate. The same extrapolation applied downwards from bC2n yields an age of 2.106 ± 0.056 Ma for the base of the section. These calculated ages are greater than the maximum age defined for the base of section in Area 131 (2.08 ± 0.04 Ma; Table 1), suggesting that accumulation rates in the lower part of the section were slightly higher, at least 23.0 ± 9 cm/kyr.

## Age of the sediments and fossils

### Age of marker beds

Our results show that intraformational conglomerate layers GPC (also named marker bed g) and LCC (marker bed h), previously assumed to be reliable marker beds for correlation between areas

(Feibel et al., 1989), are not synchronous in Areas 105 and 131: the conglomerate bed identified as 'GPC-' in Area 105 is time equivalent to conglomerate bed LCC in Area 131 (Fig. 4e and i). This demonstrates that these similar looking conglomerate beds, while useful as marker beds for local correlations (if the beds can be continuously traced between the sections), are not suitable for making regional stratigraphical correlations. For instance in Areas 130 and 115, a conglomerate bed 'GPC-' has been used as a stratigraphic marker for hominin fossils occurrences (Feibel et al., 1989; Leakey et al., 2012) but our data suggest it cannot be assumed that these conglomerate beds are synchronous with the GPC bed in Areas 131 or 105. Hence, care should be taken in making inferences on correlation and ages of these fossils from different areas, based solely upon sedimentological comparisons.

### Age control on hominin fossils

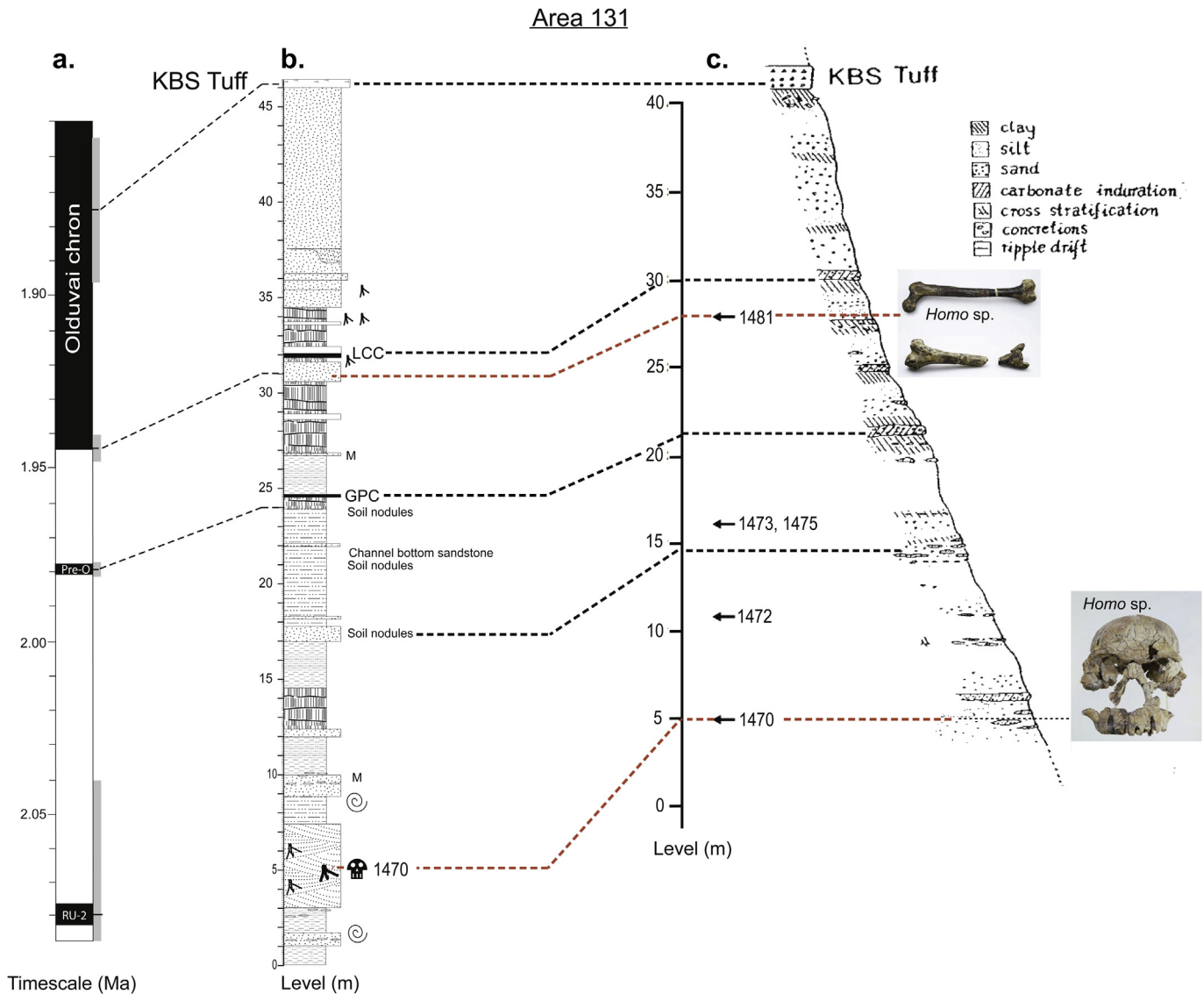
The lithology of the sampled section of Area 131 is now logged in such detail that we can correlate it with a published detailed log section providing the stratigraphic positions of hominin fossils that have been found in the UBU of this area (Leakey, 1973a,b, 1974; Leakey et al., 2012). Fig. 5 shows the lithostratigraphic correlation between the Area 131 log from this study (Fig. 5a) and the Area 131 log from Leakey (1973a; Fig. 5b). The correlation allows placement of KNM-ER 1481 in the sandy interval below GPC, between the 30.6 and 31.0 m level of our stratigraphy (Fig. 5a, see also Fig. 4e). KNM-ER 1470 can be precisely placed at the 5 m level in our stratigraphy (Figs. 4e and 5a) because our sampling section includes the site of KNM-ER 1470, and can also be correlated to the Leakey log (Fig. 5b). The other hominin fossils from the Leakey log (KNM-ER 1472, 1473, 1475) are constrained by the GPC (m level 24.6) and the stratigraphic position of KNM-ER 1470 (m level 6; Fig. 5b). KNM-ER 62000 was recently discovered in Area 131 and derives from just above the GPC bed (Leakey et al., 2012) so it can be assigned to meter level 25 in our stratigraphy (Fig. 4e). The mandible KNM-ER 1802 (Leakey, 1974) derives from below level GPC ('GPC-'; Feibel et al., 1989), and while it is not exactly known how far below, it is unambiguously older than KNM-ER 62000. Mandible KNM-ER 1482 (Leakey, 1973b) derives from a stratigraphic level between that of KNM-ER 1475 and KNM-ER 1472 (Feibel et al., 1989; Fig. 5b). Mandible KNM-ER 1801 (Leakey, 1974) was found at a level below a marker bed called F, which according to Feibel et al. (1989) is situated ~30 m below KBS Tuff, so it is stratigraphically close to KNM-ER 1472 (Fig. 5b).

It is now possible to provide improved age constraints for all 15 hominin fossils that were found in Area 131 (Leakey, 1973a,b, 1974; Feibel et al., 1989; Leakey et al., 2012; Table 7). The age of five of the hominin fossils can be more precisely controlled. KNM-ER 1481, KNM-ER 1803 and KNM-ER 1469 derive from just below the base of the Olduvai chron and ~7 m above the Pre-Olduvai event and thus are constrained between 1.945 ± 0.004 and 1.977 ± 0.002 Ma. KNM-ER 62000 derives from just above the Pre-Olduvai event and is similarly constrained between 1.945 ± 0.004 and

**Table 6**  
Calculated sediment accumulation rates between dated markers in Areas 105 and 131.

Marker	Reference	Area 105		Area 131	
	Age (ka) ± Δage (kyr)	Level (cm) ± Δlevel (cm)	Rate (cm/kyr) ± Δrate (cm/kyr)	Level (cm) ± Δlevel (cm)	Rate (cm/kyr) ± Δrate (cm/kyr)
KBS	1876 ± 21	3100.0 ± 0.0	<b>31.3 ± 9.2</b>	4600 ± 0	<b>21.6 ± 7.2</b>
bC2n	1945 ± 4	940.0 ± 30.0		3110 ± 90	<b>21.7 ± 5.1</b>
Pre-Old	1977 ± 2	<b>-61.7 ± 116.9</b>		2415 ± 45	<b>23.0 ± 8.6</b>
RU-2	2080 ± 40			0 ± 0	

Age, reference age: KBS—KBS Tuff (dated by McDougall and Brown, 2006; age normalized according to Kuiper et al., 2008); bC2n, base of Olduvai chron (Hornig et al., 2002; Lourens et al., 2004); Pre-Old, Pre-Olduvai event (Channell and Guyodo, 2004; Pickering et al., 2011). Δage, Analytical error in the age; Level, Stratigraphic position uncertainty (in bold italics if linearly extrapolated from rate at overlying tie point); Δlevel, Stratigraphic position uncertainty (in bold italics if linearly extrapolated from Δrate at overlying tie point); Rate, Sediment accumulation rate derived from linear interpolation between age tie points; Δrate, Rate uncertainty including level and age uncertainties.



**Figure 5.** Lithostratigraphic correlation of Area 131 sections. a. Timescale with periods of normal polarity in black and reversed polarity in white. Pre-O: Pre-Olduvai event. RU-2: second Réunion event. In grey: error bars on the ages of KBS Tuff and paleomagnetic reversals (Table 1). b. Composite section in this study. c. Composite section in Leakey (1973a). See legend of Table 4.

1.977 ± 0.002 Ma. We have established that KNM-ER 1470 derives from 26 m below bc2n, 19 m below the base of the Pre-Olduvai event, and 5 m above the base of the section (reported here). We can provide an age estimate for this important fossil by calculating an age based on meters below bc2n, using a calculated average accumulation rate below bc2n in Area 131. The average accumulation rate, based on a maximum age of 2.08 ± 0.04 Ma for the base of the section, is 23.0 ± 9 cm/kyr (Table 6). Our calculation results in an age of 2.058 ± 0.034 Ma for KNM-ER 1470 (Table 6). This is somewhat older than a previous age estimate for KNM-ER 1470 (2.03 ± 0.05 Ma; McDougall et al., 2012) and calculated at higher precision.

## Discussion and conclusions

### Implications of improved age control for early Homo

Several of the hominin fossils for which improved age estimates are reported here are key specimens in a recent study that

reconsiders early *Homo* in eastern Africa on the basis of new fossil evidence from Koobi Fora (Leakey et al., 2012). One of the new fossils described in that study is a partial face, KNM-ER 62000, which shows strong similarities to KNM-ER 1470. Its well-preserved palate and teeth make it possible to infer aspects of the mandible that occluded with the distinctive upper dental arcades of KNM-ER 1470 and KNM-ER 62000. Good matches are not only KNM-ER 60000, a new, nearly complete early *Homo* mandible from above the KBS Tuff in Area 105, but also the partial mandibles KNM-ER 1482 and KNM-ER 1801 from UBU deposits considered here. In contrast, the mandible KNM-ER 1802, which has been most commonly associated with KNM-ER 1470, is shown to be an unlikely match (Leakey et al., 2012). As such, Leakey et al. (2012) find both cranial and mandibular evidence for two species of early *Homo*. These are often referred to as *Homo rudolfensis* for KNM-ER 1470 and associated specimens, and *Homo habilis* for other, subnasally more prognathic specimens (Wood, 1992, 2012), but Leakey et al. (2012) refrain from assigning species names because they feel that the association of the *H. habilis* type specimen OH 7 is not yet

**Table 7**

Age constraints for hominin fossils from Area 131 (Feibel et al., 1989; Wood, 1991; Wood and Leakey, 2011; Leakey et al., 2012).

KNM-ER	Part of the specimen	Taxon	Level		Constrained between (Ma)
			Feibel et al., 1989	This study	
1483	Mandibular fragments	<i>Homo</i> sp.	KBS-4m		1.85–1.95 (between KBS Tuff and base Olduvai chron)
1481	L femur, partial tibia and fibula	<i>Homo</i> sp.	KBS-12m	KBS-15m	1.95–1.98 (just (max ~30 cm) below base Olduvai chron)
1803	R mandibular body	<i>P. boisei</i>	KBS-12m	KBS-15m	1.95–1.98 (just (max ~30 cm) below base Olduvai chron)
1469	L mandibular body	<i>P. boisei</i>	KBS-13m		1.95–1.98 (between base Olduvai chron and Pre-Olduvai event)
62000	Partial cranium	<i>Homo</i> sp.		KBS-21m*	1.95–1.98 (1 m above Pre-Olduvai event)
1802	Mandibular body	<i>Homo</i> sp.	GPC-		1.98–2.09 (between Pre-Olduvai event and age KNM-ER 1470)
3951	Postcrania	<i>Homo</i> sp.	GPC-		1.98–2.09 (between Pre-Olduvai event and age KNM-ER 1470)
1473	R proximal humerus	<i>Homo</i> sp.	KBS-25m		1.98–2.09 (between Pre-Olduvai event and age KNM-ER 1470)
1475	R proximal femur	<i>Homo</i> sp.	KBS-25m		1.98–2.09 (between Pre-Olduvai event and age KNM-ER 1470)
1482	Partial mandible	<i>Homo</i> sp.	KBS-26m		1.98–2.09 (between Pre-Olduvai event and age KNM-ER 1470)
1471	R proximal tibia	<i>Homo</i> sp.	KBS-27m		1.98–2.09 (between Pre-Olduvai event and age KNM-ER 1470)
1472	R Femur	<i>Homo</i> sp.	KBS-30m		1.98–2.09 (between Pre-Olduvai event and age KNM-ER 1470)
1801	L mandibular body	<i>Homo</i> sp.	F-		1.98–2.09 (between Pre-Olduvai event and age KNM-ER 1470)
1474	Vault fragment	<i>Homo</i> sp.	KBS-32m		1.98–2.09 (between Pre-Olduvai event and age KNM-ER 1470)
1470	Cranium	<i>Homo</i> sp.	KBS-36m	KBS-41m	2.03–2.09

R: Right, L: Left.

\*NB In the composite section of Leakey et al. (2012) and in our composite section, KNM-ER 62000 derives from a level 21 m below KBS Tuff (KBS-21m). In individual section CSF131-2, KNM-ER 62000 derives from KBS-17 m (SOM; Leakey et al., 2012).

fully resolved. Our current study provides older ages for KNM-ER 1470, KNM-ER 1482, KNM-ER 1801 and KNM-ER 62000 than previously estimated (Feibel et al., 1989; Leakey et al., 2012). These data appear to support an interpretation of the early *Homo* fossil record that specimens assigned to *H. rudolfensis* predate those assigned to *H. habilis*, i.e., these species form time-successive segments of a single anagenetic evolutionary lineage rather than separate, contemporaneous species (Suwa et al., 1996). However, our findings equally increase the age of KNM-ER 1802, a specimen unlikely to be associated with KNM-ER 1470 and KNM-ER 62000. Moreover, early occurrence of a more subnasally prognathic morphology is shown by *Homo* specimens OH 24 (1.9 Ma; Hay, 1976; Walter et al., 1991) and A.L. 666-1 (2.33 Ma; Kimbel et al., 1996), whereas the occurrence of the 'KNM-ER 1470 group' in younger, post-UBU times is indicated by KNM-ER 60000 dated to 1.78–1.87 Ma (Leakey et al., 2012). Hence, the dates now available are consistent with a temporal overlap of two species of early *Homo* that can be distinguished by their facial morphology (sensu Leakey et al., 2012).

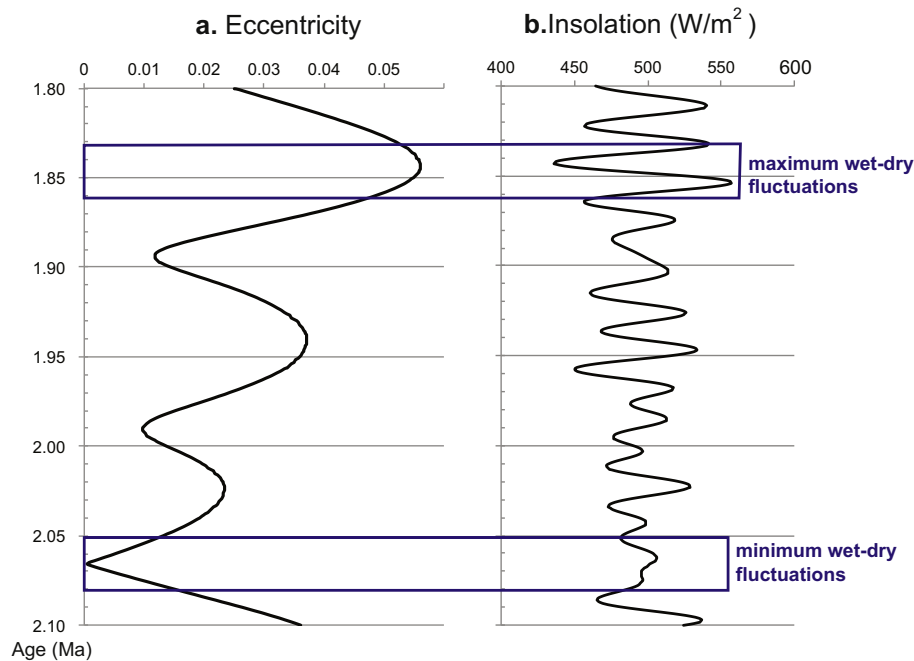
The revised, older dates obtained here for early *Homo* specimens do not support the suggestion that *Australopithecus sediba* preserved at Malapa, South Africa, could have been ancestral to *Homo* because at  $1.977 \pm 0.002$  Ma it would be older than the oldest uncontested representatives of this genus (Pickering et al., 2011). Regardless of other arguments, such as the premature dismissal of the *Homo* affinities of the A.L. 666-1 maxilla from Hadar (Spoor, 2011), it is now clear that several early *Homo* fossils reassessed here are chronologically constrained between  $1.977 \pm 0.002$  Ma and the base of section at  $2.08 \pm 0.04$  Ma, and thus older than any known *A. sediba* specimens. Most prominent among these fossils is KNM-ER 1470, now dated to  $2.058 \pm 0.034$  Ma.

The femur KNM-ER 1481 derives from just (at maximum ~30 cm; Figs. 4 and 5) under the base of the Olduvai chron, resulting in an age of slightly older than  $1.945 \pm 0.004$  (Fig. 4e). Wood and Leakey (2011) attribute this fossil to *Homo erectus*. If confirmed, this would make KNM-ER 1481 the oldest securely dated *H. erectus* fossil found so far, just older than the innominate KNM-ER 3228, also attributed to this species by Wood and Leakey (2011) and assigned an astronomically tuned age of 1.92 Ma (Joordens et al., 2011). Nevertheless, at present it cannot be excluded that the femur and innominate belong to *H. rudolfensis* instead of *H. erectus*. The prime importance of the dating evidence is that it confirms the presence of a hominin with modern human-like locomotion (Richmond and Jungers, 2008; Harmon, 2009) in the Turkana Basin as early as ~1.95 Ma.

#### Implications of orbitally forced climate cycles for hominin presence

Joordens et al. (2011) concluded that in the time interval between ~1.87 and 1.99 Ma, hominins from the Koobi Fora Ridge and Bura Hasuma in the eastern Turkana Basin were present throughout dry and wet phases of precession cycles. Moreover, this study suggested that the lacustrine Turkana Basin was an oasis-like refugium for water-dependent fauna during regionally dry periods. This hypothesis can now be tested with the well-dated hominin fossils from the Karari Ridge. Our Sr isotope and magnetostratigraphic results indicate that KNM-ER 1481 and KNM-ER 1803, deriving from just below bc2n, can be correlated with a dry phase of the precession cycle (Figs. 4 and 5, Table 7). KNM-ER 1470 occurred in a relatively wet phase of the precession cycle (Fig. 4d and e). Based on stratigraphic levels of KNM-ER 1474, KNM-ER 1801, KNM-ER 1472 and KNM-ER 1471 (deriving from levels ~4–6 and 9 m, respectively, above KNM-ER 1470; Table 7) we can infer that they are correlated with different phases of that same precession cycle (Fig. 4). Thus, our results suggest that also in the time interval between  $2.058 \pm 0.034$  and  $1.945 \pm 0.004$  Ma, hominins occurred throughout the wet–dry cycles, providing further support for the proposed hypothesis that the lacustrine Turkana Basin was an aridity refugium during regionally dry periods (Joordens et al., 2011).

However, it should be taken into account that precession cycles are modulated by the 100 kyr and 400 kyr Earth eccentricity cycles: the more oblate the Earth's orbit (higher eccentricity value), the larger is the amplitude of the precession and insolation cycles (Fig. 6; Laskar et al., 2004). Given the importance of precession forcing to the monsoon, eccentricity is expected to influence the intensity of monsoonal fluctuations (e.g., Wang et al., 2010). Hence, during an eccentricity minimum, the variation in monsoonal intensity and thus of Omo River water input over a precession cycle would be minimal, and as a result the amplitude in Sr isotope cycles would be small. During the early UBU time interval and especially in the period between 2.08 and 1.98 Ma, eccentricity values were very low, with a minimum of ~0.005 occurring at 2.06 Ma (Laskar et al., 2004; Fig. 6). As a consequence, the amplitude of monsoonal wet–dry fluctuations, and hence of fluctuations in Omo River input into the Turkana Basin, can be expected to have been minimal then. At 1.94–1.95 Ma, the eccentricity value was relatively high (~0.037; Fig. 6) and the amplitude of the fluctuations in Omo River water input would have been relatively large then. Indeed, occurrence of this phenomenon appears to be reflected by the small



**Figure 6.** Orbital cycles. a. Earth eccentricity, b. summer insolation at 65°N. We apply the 65°N summer insolation of the La2004(1,1) astronomical solution (Laskar et al., 2004) because this target curve clearly reflects the distinct precession-obliquity interference patterns recorded by the monsoon intensity related sapropel depositions in the Mediterranean (Lourens et al., 1996, 2004).

amplitude (0.0002) of the Sr isotope cycle in Area 131 around 2.06 Ma, compared with the larger amplitude (0.0006) of the younger cycle around 1.95 Ma in Area 105 (Figs. 4 and 6).

In the time interval just above the KBS Tuff, from ~1.87 to 1.82 Ma, the eccentricity was maximal (Fig. 6), which would have caused more extreme fluctuations in precessional (monsoonal intensity) amplitude, and hence in Omo River input into the lake. This is corroborated by the observed major lake level fluctuations in the KBS Member overlying the UBU (Lepre et al., 2007). It is therefore recommended to test the aridity refuge hypothesis also in deposits above the KBS Tuff in order to study if/how the high eccentricity value affected hominin presence in the fluctuating lakeshore deposits. The improved age control presented in this study can also be applied to non-hominin fossil faunal occurrences in the upper Burgi Member, allowing future analysis of the sensitivity of faunal communities to eccentricity-modulated precessional patterns of climatic and environmental change.

Finally, we wish to draw attention to the implications of our results in a wider biogeographical context. By providing an age estimate for the level at which KNM-ER 1470 was found, we also establish the (observed) first appearance datum of the stingray *Dasyatis africana* in the Turkana Basin at  $2.058 \pm 0.034$  Ma (Fig. 4e), instead of at ~1.90 Ma as previously estimated (Feibel, 1993). The sudden appearance of marine-derived stingrays in the fossil record (combined with the absence of stingray fossils from older levels) indicates that the appearance signifies the establishment of a hydrographic connection, and hence a possible dispersal route for water-dependent fauna, between the Indian Ocean and the Turkana Basin. Leakey et al. (2008) concluded that many species of fossil cercopithecoid monkeys that first appeared in the Turkana Basin without obvious local predecessors are likely to have dispersed to the basin while having evolved elsewhere. The authors inferred that waterways with gallery forests linking the Turkana Basin with habitats elsewhere would have been critical in facilitating this dispersal, and recognized that Central African forests and particularly the eastern African coastal forest along the Indian Ocean would have provided ideal habitats for primate

evolution (as proposed by Kingdon, 2003). This could have also been the case for hominins (Kingdon, 2003; Joordens et al., 2012), and we therefore propose that the Indian Ocean coastal strip should be considered as a possible evolutionary source area for one or more of the multiple *Homo* species that appeared in the Turkana Basin during the crucial time interval around 2 Ma.

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#### Appendix A. Supplementary data

Supplementary data related to this article can be found at <http://dx.doi.org/10.1016/j.jhevol.2013.09.002>.

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